



# MAVIS

**DEEPER THAN HST, SHARPER THAN JWST**  
**4 SLIDES STATUS @ START OF PHASE B**

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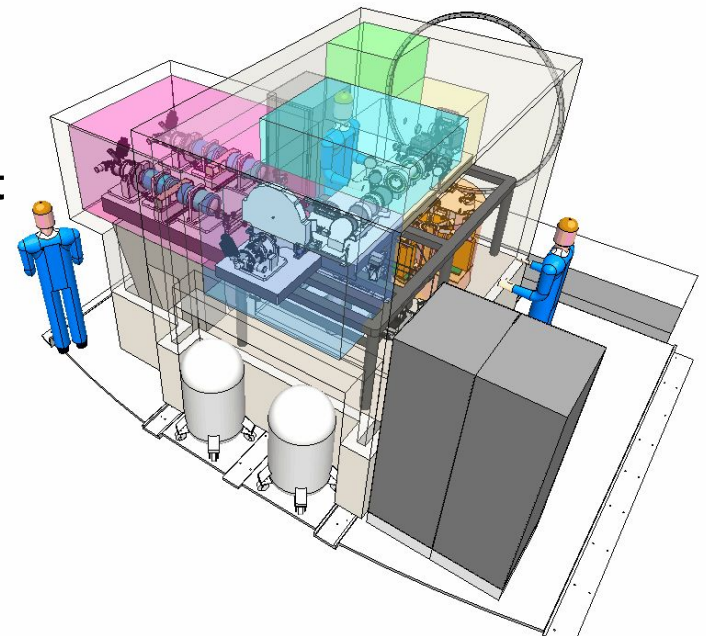
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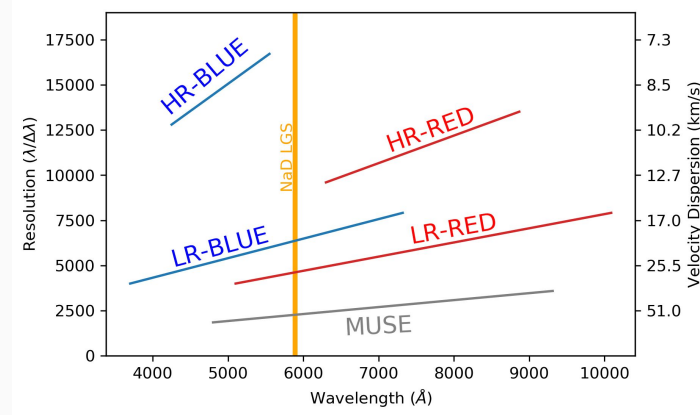
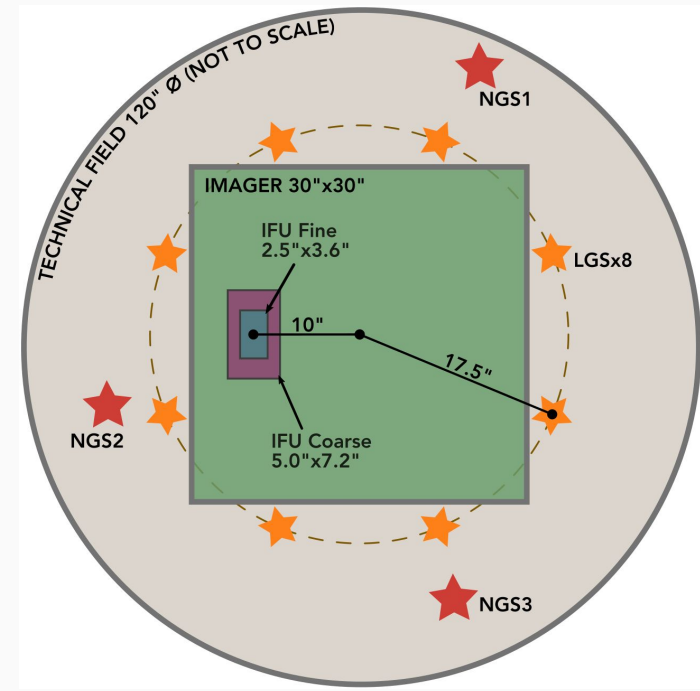
# MAVIS: DEEPER THAN HST, SHARPER THAN JWST

- **Multi-conjugate Adaptive Optics** system for correction in the **visible**
  - complete with an **4kX4k imager** @ 7.3mas pixels
  - and an **IFU w/ 4 spectral resolution modes** (4-12k), ¼ number of MUSE spaxels
- Expecting > **10% Strehl** (goal 15%) at **V band** over **30"x30"**
- Consortium **Australia (AAO Consortium, lead) / INAF / LAM / ESO**
- Passed phase A 06/2020, **first light expected Sem 2 2027**
- For the ESO VLT AOF (UT4)
  - 4x2 Laser Guide Stars;
  - 3 Near-IR NGS Wavefront Sensors (using SAPHIRA);
  - 3 Deformable mirrors (DSM + 2 post focal DMs);
- A **brilliant science case** for a **facility instrument** ([publicly available on arXiv](#)), with themes:
  - Emergence of Hubble sequence
  - Resolving galaxy contents
  - Star clusters as tracers of galaxy evolution
  - Birth, life, death of stars and planets



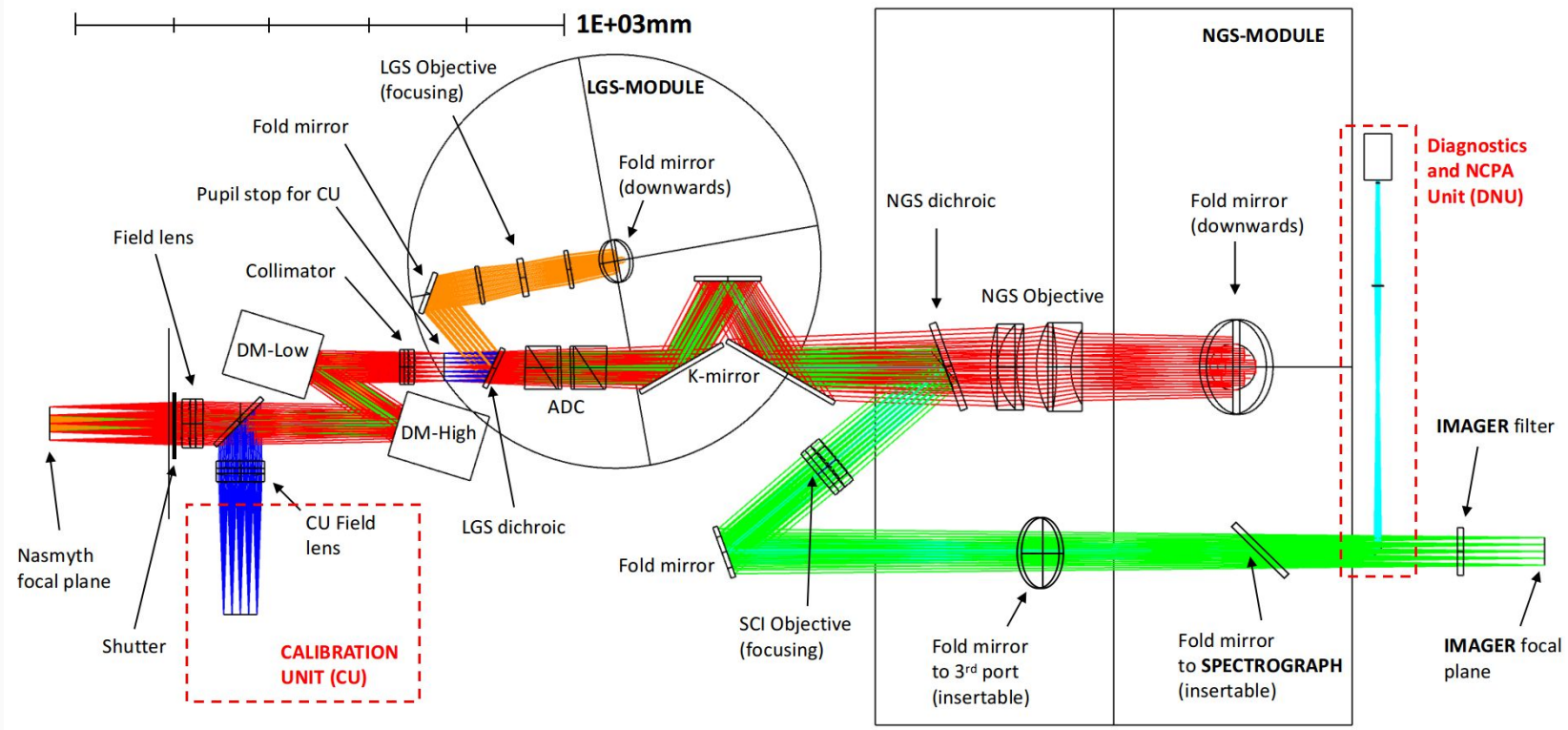
# MAVIS TOP-LEVEL SPECIFICATIONS

<b>Science Field</b>	30"x30"
<b>Angular Resolution</b>	FWHM ~ 20mas at V band
<b>Strehl Ratio</b>	>10% (15% goal) at V under median conditions
<b>Sky Coverage</b>	> 50% at the Galactic pole, 15% EE in 50mas spaxel
<b>Wavelength Coverage</b>	VRI (optimised); B-z (extended)
<b>Imager</b>	~ 7mas pixels. 7 broad and 15 narrow band filters, 1h 10 $\sigma$ for V ~ 29.5 (HST SNR x2)
<b>Spectrograph</b>	Image slicer. Two spatial modes: ~3"x3" @ 25mas and ~6"x6" @ 50mas. Four spectral modes: 370-1000nm, R=5,000-15,000
<b>Visitor port</b>	Potential for third instrument



# AN INNOVATIVE OPTICAL DESIGN

- Uses refractive elements in main path - essentially no field distortions
- No collimated beams
- Everything gravity invariant (except K mirror and ADC)
- Very “healthy” design, all zero-order principles of AO design are there.



## THE MAVIS CONSORTIUM

- Australia/INAF/LAM/ESO
- First ESO instrument led by Australia



AAO-INAF-LAM-ESO